

# M-Theory on Calabi-Yau 5-folds

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Work with Alexander Haupt and

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previous work with Hong Lü, Chris Pope and

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# Motivation

- ◆  $CY_3$  manifolds provide one of the most important approaches to phenomenological contact between realistic physics and string/M-theory.
- ◆ The standard embedding of an  $SU(3)$  spin connection into the heterotic string's  $E_8 \times E_8$  gauge group breaks the YM gauge group down to  $E_8 \times E_6$  and  $E_6$  is physically appealing.
- ◆ At the same time, from an M-theory perspective, the 4+7 split is unnatural. A more “democratic” formulation of the spatial dimensions would seem more natural.
- ◆ Cosmology could naturally involve a 1+10 split. All space dimensions would initially be treated as compact, in anticipation of 3 of them expanding.

# Overview

- ◆ Review of bosonic sector of  $D=11$  supergravity including normalizations Bilal
- ◆ Topological considerations and flux quantization in M-theory
  - topological constraint on compact 10-manifolds
- ◆ CY moduli sigma model
- ◆ 2-component local supersymmetry in  $D=11$
- ◆ Effect of  $\alpha'$  corrections on  $CY_5$  geometrical structure
- ◆ Supersymmetry preservation and generalized holonomy

# D=11 supergravity

$$I_{11} = I_{CJS,B} + I_{CJS,F} + I_{GS} + \dots$$

$$I_{CJS,B} = \frac{1}{2\kappa_{11}^2} \int_{\mathcal{M}} \left\{ R * 1 - \frac{1}{2} G \wedge *G - \frac{1}{6} G \wedge G \wedge C \right\}$$

$$G_{[4]} = dC_{[3]}$$

4-form field strength for  
3-form gauge field

$$I_{CJS,F} = -\frac{1}{2\kappa_{11}^2} \int_{\mathcal{M}} d^{11}x \sqrt{-g} \left\{ \bar{\psi}_M \Gamma^{MNP} D_N(\omega) \psi_P \right. \\ \left. + \frac{1}{96} (\bar{\psi}_M \Gamma^{MNPQRS} \psi_S + 12 \bar{\psi}^N \Gamma^{PQ} \psi^R) G_{NPQR} + (\text{fermi})^4 \right\}$$

- ◆ The above terms combine to form an invariant under the classical supersymmetry transformations

$$\delta_\epsilon g_{MN} = 2\bar{\epsilon} \Gamma_{(M} \psi_{N)},$$

$$\delta_\epsilon C_{MNP} = -3\bar{\epsilon} \Gamma_{[MN} \psi_{P]},$$

$$\delta_\epsilon \psi_M = 2D_M(\omega)\epsilon + \frac{1}{144} (\Gamma_M^{NPQR} - 8\delta_M^N \Gamma^{PQR}) \epsilon G_{NPQR} + (\text{fermi})^3.$$

- ◆ Variation of the Cremmer-Julia-Scherk action leads to the classical supergravity field equations:

$$R_{MN} = \frac{1}{12} G_{MM_2\dots M_4} G_N{}^{M_2\dots M_4} - \frac{1}{144} g_{MN} G_{M_1\dots M_4} G^{M_1\dots M_4}$$

$$d * G + \frac{1}{2} G \wedge G = 0.$$

$$\Gamma^{MNP} D_N(\omega) \psi_P + \frac{1}{96} (\Gamma^{MNPQRS} \psi_S + 12 \delta^{MN} \Gamma^{PQ} \psi^R) G_{NPQR} + (\text{fermi})^3 = 0$$

- ◆ Quantum corrections change these equations in a way that is important for CY<sub>5</sub> compactifications. Among the  $\beta = (2\pi)^2 \alpha'^3$  quantum corrections is a Green-Schwarz type term needed for M<sub>5</sub>-brane worldvolume anomaly cancellations.

Vafa & Witten

Duff, Liu & Minasian

- ◆ This GS term is a superpartner of the  $R^4_{\mu\nu\rho\sigma}$  effective action corrections.

- ◆ The classical CJS equation for  $C_{[3]}$

$$d * G + \frac{1}{2} G \wedge G = 0$$

is accordingly modified by the Green-Schwarz correction

$$I_{GS} = \frac{-(2\pi)^4 \beta}{2\kappa_{11}^2} \int C \wedge X_8$$

where 
$$X_8 = \frac{1}{(2\pi)^4} \left[ -\frac{1}{768} (tr R^2)^2 + \frac{1}{192} tr R^4 \right]$$

- ◆ This gives rise to the quantum-corrected equation

$$d * G + \frac{1}{2} G \wedge G + (2\pi)^4 \beta X_8 = 0.$$

- ◆ The Green-Schwarz correction term is necessary for cancelation of anomalies on the  $d=6$  worldvolumes of 5-branes:  $\beta = \frac{1}{(2\pi)^3 T_5}$   $T_5 = 5\text{-brane tension}$

- ◆ One also has the Dirac quantization condition  $T_2 T_5 = \frac{2\pi}{2\kappa_{11}^2}$   $T_2 = 2\text{-brane tension}$  and the condition

$$T_5 = \frac{1}{2\pi} T_2^2 \quad \text{de Alwis}$$

which is needed, e.g., for invariance under large 3-form gauge transformations.

Lavrinenko, Lü, Pope & K.S.S  
Kalkkinen & K.S.S

- ◆ Putting these together, have

$$T_2 = \left( \frac{2\pi^2}{\kappa_{11}^2} \right)^{1/3} \quad \beta = \left( \frac{2\kappa_{11}^2}{(2\pi)^5} \right)^{2/3} \quad 2\kappa_{11}^2 = (2\pi)^8 (\alpha')^{9/2}$$

# Topological considerations

A. Haupt, A. Lukas & K.S.S.

- ◆ Corrected 3-form field equation:

$$d * G + \frac{1}{2} G \wedge G + (2\pi)^4 \beta X_8 = 0$$

where

$$X_8 = \frac{1}{48} \left( \left( \frac{p_1}{2} \right)^2 - p_2 \right)$$

$$p_1 = -\frac{1}{2} \left( \frac{1}{2\pi} \right)^2 \text{tr} R^2$$

$$p_2 = \frac{1}{8} \left( \frac{1}{2\pi} \right)^4 \left( (\text{tr} R^2)^2 - 2 \text{tr} R^4 \right)$$

1st & 2nd  
Pontriagin classes

- ◆ Now specialize to  $M_{11} = \mathbb{R} \times CY_5$  and simplify above relations:

$$p(T(\mathbb{R} \times CY_5)) = p(T(\mathbb{R})) \wedge p(T(CY_5))$$

$$p(T(\mathbb{R})) = 1 \text{ so } p(T(M_{10})) \text{ is given by } p(T(CY_5))$$



- ◆ Now, for complex manifolds, there are relations between Pontriagin and Chern classes:

$$p_1 = c_1^2 - 2c_2$$

T. Hübsch

$$p_2 = 2c_4 - 2c_1c_3 + c_2^2$$

so for the case of a Calabi-Yau manifold with

$c_1 = 0$  one has  $\left(\frac{p_1}{2}\right)^2 - p_2 = -2c_4$

and consequently  $X_8 = -\frac{1}{24}c_4$

- ◆ Define  $g = \frac{1}{(2\pi)^2\beta^{1/2}}G$  and use the corrected field equations together with the fact that  $d * G$  is exact to deduce  $\left[\frac{1}{2}G \wedge G + (2\pi)^4\beta X_8\right] = 0$  giving the topological constraint

$$c_4(CY_5) - 12[g] \wedge [g] = 0$$

# 4-form flux quantization

- ◆ 2-branes couple to the  $C_{[3]}$  background via

$$S_{WZ}^{2br} = T_2 \int_{W_3} C \rightarrow T_2 \int_{D_4} G \quad \partial D_4 = W_3$$

- ◆ This gives the flux quantization condition

$$[g] - \frac{p_1}{4} \in H^4(CY_5, \mathbb{Z}) \quad \text{Witten}$$

or, for  $c_1 = 0$ ,

$$[g] + \frac{c_2}{2} \in H^4(CY_5, \mathbb{Z}) \quad g = \frac{T_2}{2\pi} G$$

- ◆ Thus, depending on the value of the 2<sup>nd</sup> Chern class  $c_2$ , the normalized flux  $g$  is quantized in integer or half-integer units.

- ◆ Happily, this is consistent with the topological constraint  $c_4(CY_5) - 12[g] \wedge [g] = 0$

- ◆ For complete intersection compact  $CY_5$ , analysis shows that  $c_4(CY_5^{c.i.}) > 0$  requiring  $[g] \neq 0$  so 4-form flux must be turned on at order  $\sqrt{\beta}$
- ◆ However, one can make orbifold constructions with  $c_4 = 0$ .
- ◆ Non-compact  $CY_5$  can also have  $c_4 = 0$ .
- ◆ In cases with  $c_4 = 0$ , the flux is turned on at order  $\beta$

# CY<sub>5</sub> moduli $D = 1$ Supersymmetric sigma model

- ◆ CY<sub>5</sub> Hodge diamond:

$$\begin{array}{cccccc} & & & & & 1 \\ & & & & 0 & 0 \\ & & & 0 & h^{1,1} & 0 \\ & & 0 & h^{1,2} & h^{1,2} & 0 \\ & 0 & h^{1,3} & h^{2,2} & h^{1,3} & 0 \\ 1 & h^{1,4} & h^{2,3} & h^{2,3} & h^{1,4} & 1 \end{array}$$

- ◆ Hirzebruch-Riemann-Roch theorem with  $c_1 = 0$ :

$$11h^{1,1} - 10h^{1,2} - h^{2,2} + h^{2,3} + 10h^{1,3} - 11h^{1,4} = 0$$

so there are  $6-1=5$  independent Hodge numbers. The corresponding harmonic forms contribute  $D = 1$  massless Kaluza-Klein modes.

◆ Metric:

$$ds^2 = -N d\tau^2 + 2g_{rs}(x, \varphi^I(\tau)) dx^r dx^s$$

$$\varphi^I(\tau) = (t^i(\tau), z^a(\tau), z^{\bar{a}}(\tau))$$

$h^{1,1}$   $h^{1,4}$  moduli

in complex coordinates

$$x^r \rightarrow x^\mu, x^{\bar{\nu}} \quad \mu, \bar{\nu} = 1, \dots, 5$$

$$\delta g_{\mu\bar{\nu}} = \delta t^i \omega_{i\mu\bar{\nu}} \quad \delta g_{\mu\nu} = \delta z^{\bar{a}} b_{\bar{a}\mu\nu} \quad \delta g_{\bar{\mu}\bar{\nu}} = \delta z^a b_{a\bar{\mu}\bar{\nu}}$$

$$b_{\bar{a}\mu\nu} = \frac{i}{\|\Omega\|^2} \Omega_\mu^{\bar{\rho}\bar{\sigma}\bar{\tau}} \chi_{\bar{a}\bar{\rho}\bar{\sigma}\bar{\tau}\nu}$$

(5,0) volume form

(4,1) harmonic form

$$\omega_i \in \text{Harm}(1,1)$$

$$\chi_a \in \text{Harm}(1,4)$$

◆ 3-form field:

$$\delta C = \xi^p(\tau) \mathbf{v}_p + \text{c.c.}$$

$h^{1,2}$

$$\mathbf{v}_p \in \text{Harm}(1,2)$$

# Fermionic zero modes

- ◆ Expand  $\Psi_M(\tau, x^r)$  using the Killing spinor  $\eta(x^r)$  on  $CY_5$ ,

$$\text{e.g. } \Psi_0(\tau, x^r) = \bar{\psi}_0(\tau)\eta(x^r) + \text{cc} \quad \eta^\dagger\eta = 1$$

- ◆ For  $\Psi_\mu(x^r)$ ,  $\Psi_{\bar{\nu}}(x^r)$  the expansion uses the

(1,1), (2,1), (3,1) and (4,1) harmonic forms:

$$\begin{aligned} \psi_{\bar{\mu}} = & \psi^i(\tau) \otimes \overset{(1,1)}{(\omega_{i\alpha_1\bar{\mu}}\gamma^{\alpha_1}\eta)} + \frac{1}{4}\lambda^p(\tau) \otimes \overset{(2,1)}{(\nu_{p\alpha_1\alpha_2\bar{\mu}}\gamma^{\alpha_1\alpha_2}\eta)} \\ & + \frac{1}{4!}\rho^x(\tau) \otimes \overset{(3,1)}{(\varpi_{x\alpha_1\dots\alpha_3\bar{\mu}}\gamma^{\alpha_1\dots\alpha_3}\eta)} - \frac{1}{4!}\kappa^a(\tau) \otimes \overset{(4,1)}{(\|\Omega\|^{-1}\chi_{a\alpha_1\dots\alpha_4\bar{\mu}}\gamma^{\alpha_1\dots\alpha_4}\eta)} \\ \psi_\mu = & (\psi_{\bar{\mu}})^* \end{aligned}$$

- ◆ The (3,1) species has no bosonic partners, however. This points out a strange feature of supersymmetric life in  $D = 1$ : on-shell bosonic and fermionic degrees of freedom do not have to balance.

Coles &  
Papadopoulos

- ◆ What happens to the other possible types of harmonic forms, e.g. (3,2), (2,2) and (5,0)?
  - These are reabsorbed into the (1,1) and (2,1) harmonic types.
  - To see this, one needs to use the  $\gamma^{\bar{u}}\eta = 0$  property of CY Killing spinors together with the Dirac algebra  $\{\gamma^{\mu}, \gamma^{\bar{\nu}}\} = 2g^{\mu\bar{\nu}}$  and Fierz identities to reduce these species to other types. E.g. the (5,0) type is converted into a (1,1) species, and is the superpartner of the CY volume modulus.

# Bosonic sigma model

gauge N=1

$$I_{\text{CJS}}^{\text{B}} \xrightarrow{M_{11}=\mathbb{R}\times\text{CY}_5} \int d\tau \left\{ \frac{1}{4} G_{ij}^{(1,1)}(t) \dot{t}^i \dot{t}^j + G_{p\bar{q}}^{(2,1)}(t) \dot{\xi}^p \dot{\bar{\xi}}^{\bar{q}} - 4V(t) G_{a\bar{b}}^{(4,1)}(z, \bar{z}) \dot{z}^a \dot{\bar{z}}^{\bar{b}} \right\}$$

$$G_{ij}^{(1,1)} = \partial_i \partial_j K^{(1,1)} - 25 \frac{K_i K_j}{K^2} \quad K^{(1,1)} = -\frac{1}{2} \ln K$$

$$K = \int J \wedge J \wedge J \wedge J \wedge J \quad J = t^i \omega_i \quad \text{complex structure}$$

$$= d_{i_1 \dots i_5} t^{i_1} \dots t^{i_5} \quad d_{i_1 \dots i_5}: \text{intersection numbers}$$

$$K_i = \int \omega_i \wedge J \wedge J \wedge J \wedge J = d_{ij_1 \dots j_4} t^{j_1 \dots j_4}$$

$$G_{a\bar{b}}^{(4,1)} = \partial_a \partial_{\bar{b}} K^{(4,1)} \quad K^{(4,1)} = -\ln(i(G_{\bar{a}} z^{\bar{a}} - z^a \bar{G}_a))$$

$$G_{p\bar{q}}^{(2,1)} = -2 \int_X \mathbf{v}_p \wedge \ast \bar{\mathbf{v}}_{\bar{q}} = i d_{p\bar{q}ij} t^i t^j \quad \text{Canonical inner product}$$



## ◆ Notes

- The  $(1,1)$  metric is not a canonical special Kähler metric but it is determined by intersection numbers (topological data), as is the canonical  $(2,1)$  metric.
- The  $(4,1)$  metric is the canonical Weil-Peterson metric (very special Kähler) but it is determined by a prepotential (involving non-topological data).
- The Kähler and complex structure sectors don't decouple owing to the  $V(t)$  factor.

# D=1 supersymmetry multiplets

- ◆ Inserting the D=11 supersymmetry transformations into the reduction ansatz, one finds the surviving 2-component D=1 supersymmetry (CY<sub>5</sub> breaks supersymmetry to 1/16).

- ◆ One finds two kinds of D=1 supermultiplets

- (2a) real  $\phi = \bar{\phi}$        $\phi = \varphi + i\theta\psi + i\bar{\theta}\bar{\psi} - \frac{1}{2}\theta\bar{\theta}f$

- (2b) *i.e.* (2,0) chiral  $\bar{D}Z = 0$        $Z = z + \theta\kappa - \frac{i}{2}\theta\bar{\theta}\dot{z}$

- ◆ Local D=1 supersymmetry is described by the

supervielbeins  $E_M^A$ ,  $\nabla_A = E_A^M \partial_M$ ,  $[\nabla_A, \nabla_B] = -T_{AB}^C \nabla_C$

subject to the torsion constraints

No D=1 curvature!

$$T_{\underline{\theta}\bar{\theta}}^0 = i \quad (0), \quad T_{\underline{\theta}\bar{\theta}}^\theta = 0 \quad \left(\frac{1}{2}\right) \quad \text{“conventional”}$$

$$T_{\bar{\theta}\bar{\theta}}^0 = 0 \quad (0), \quad T_{\bar{\theta}\bar{\theta}}^\theta = 0 \quad \left(\frac{1}{2}\right) \quad \text{“representation preserving”}$$

$$T_{\underline{\theta}\underline{\theta}}^\theta = 0 \quad \left(\frac{1}{2}\right) \quad \text{“type 3”}$$

- ◆ D=1 supergravity plays an entirely destructive rôle: it's effect is merely to impose constraints on the D=1 supermatter that couples to it. Subject to the torsion constraints, the remaining supergravity fields are the einbein and D=1 gravitino, contained in

$$\mathcal{E} := \text{sdet} E_A^B = N - \frac{i}{2} \theta \bar{\psi}_0 - \frac{i}{2} \bar{\theta} \psi_0$$

- Consider for example a supergravity coupled (2b) action for a single multiplet  $S = \int d\tau d^2\theta \mathcal{E} \nabla Z \bar{\nabla} \bar{Z} = \int d\tau \mathcal{L}$ .

In component fields, this Lagrangian is

$$\mathcal{L} = N^{-1} \dot{Z} \dot{\bar{Z}} - \frac{i}{2} (\kappa \dot{\bar{\kappa}} - \dot{\kappa} \bar{\kappa}) - N^{-1} (\psi_0 \kappa \dot{Z} + \bar{\psi}_0 \bar{\kappa} \dot{Z}) - N^{-1} \psi_0 \bar{\psi}_0 \kappa \bar{\kappa}$$

and varying with respect to  $N$  and  $\Psi_0$  one finds

$$Z = (\text{const.}) \quad \kappa = 0$$

- ◆ In the full supergravity-coupled action, the constraints link the (2a) and (2b) sectors.

- ◆ The full D=1 supergravity-coupled action is

$$I_1 = I_{11} \Big|_{\mathbb{R} \times X} = -\frac{m}{2} \int d\tau d^2\theta \mathcal{E} \left\{ G_{ij}^{(1,1)}(T) \nabla T^i \bar{\nabla} T^j + G_{p\bar{q}}^{(2,1)}(T) \nabla \Xi^p \bar{\nabla} \bar{\Xi}^{\bar{q}} \right. \\ \left. + G_{x\bar{y}}^{(3,1)}(T) \hat{\mathcal{R}}^x \bar{\hat{\mathcal{R}}}^{\bar{y}} + 4\mathcal{V}(T) G_{a\bar{b}}^{(4,1)}(Z, \bar{Z}) \bar{\nabla} \mathcal{Z}^a \nabla \bar{\mathcal{Z}}^{\bar{b}} \right\}$$

- ◆ Agreement between this superspace action and the Kaluza-Klein dimensionally reduced action has been checked through  $(\text{fermi})^2$  terms. The leading bosonic terms reproduce the component action given above.
- ◆ *After varying the action to obtain the supergravity constraints, one can make the gauge choices*

$$N = 1, \quad \psi_0 = 0$$

# Quantum $\beta \leftrightarrow \alpha'^3$ corrections

Lü, Pope,  
Townsend, K.S.S

- ◆ The  $\beta \int C_{[3]} \wedge X_8$  term is a D=11 superpartner to other bosonic corrections including  $R_{ABCD}^4$  terms.
- ◆ Specialize to the topologically simplest case where  $c_4 = 0$  - either noncompact  $CY_5$  or an orbifold construction.

- ◆ Correction terms of relevance:

$$\Delta L = \frac{\beta}{1152} (Y + 2Y_2 + \dots)^* + (2\pi)^4 \beta C \wedge X_8$$

plus terms that vanish for  $R_{MN} = 0$

Gross & Witten;  
Peeters, Vanhove &  
Westerberg

$$Y_{\text{string light cone}} \sim \int d^{16} \psi \exp [(\bar{\psi}_- \Gamma^{ij} \psi_-)(\bar{\psi}_+ \Gamma^{kl} \psi_+) R_{ijkl}]$$

Indices  
extended  
to 11 values

- D=11 extension of type IIA string correction
- Berezin integral  $\rightarrow R^4$  terms only

- ◆ Varying  $Y$ , get for initially Ricci-flat spaces

$$\delta \int \sqrt{-g} Y d^{11}x = \int \sqrt{-g} (X_{rs} + \nabla_r \nabla_s Z - g_{rs} \square Z) \delta g^{rs}$$

$$X_{rs} = \nabla^t \nabla^u X_{rstu}$$

$X_{rstu}$  cubic in curvatures

$$Z = R_{ijkl} R^{klmn} R_{mn}{}^{ij} - 2R_{ikjl} R^{kmln} R_m{}^i{}_n{}^j$$

- ◆ The  $Y_2$  correction term is of Lovelock form:

$$Y_2 = \frac{315}{2} R^{[m_1 m_2}{}_{m_1 m_2} \cdots R^{m_7 m_8]}{}_{m_7 m_8}$$

Lift to  $D=11$  of  
 $D=8$  Euler integrand

- ◆ Varying  $Y_2$ , get

$$\delta \int \sqrt{-g} Y_2 = \int \sqrt{-g} E_{mn} \delta g^{mn}$$

Lovelock  
Deruelle

$$E_m{}^n = -\frac{9!}{2^9} \delta_{mm_1 \cdots m_8}^{nn_1 \cdots n_8} R^{m_1 m_2}{}_{n_1 n_2} \cdots R^{m_7 m_8}{}_{n_7 n_8}$$

- ◆ Consequently, the corrected field equations are

$$\hat{R}_{00} - \frac{1}{2}g_{00}\hat{R} = -\frac{\beta}{1152}\square Z g_{00} + \frac{\beta}{576}E_{00} \quad \hat{R}_{mn}: D=11 \text{ Ricci}$$

$$\hat{R}_{ij} - \frac{1}{2}g_{ij}\hat{R} = \frac{\beta}{1152}(X_{ij} + \nabla_i \nabla_j Z - g_{ij}Z) + \frac{\beta}{576}E_{ij}$$

- ◆ To solve these, we need to introduce a *warp factor* in the metric:

$$ds_{11}^2 = -e^{2A(x^r)} d\tau^2 + e^{-\frac{1}{4}A(x^r)} ds_{10}^2$$

- then for the Ricci tensor one has

$$\hat{R}_{00} = \square A \quad \hat{R}_{ij} = R_{ij} + \frac{1}{8}g_{ij}\square A$$

$R_{ij}$ : D=10 Ricci

$\square = \nabla^2$

so  $\hat{R} = R + \frac{1}{4}\square A$  and hence

$$R_{ij} = \frac{\beta}{1152} \left( X_{ij} + \nabla_i \nabla_j Z + 2E_{ij} - \frac{1}{4}E_k^k g_{ij} \right)$$

- ◆ For an initially Kähler manifold, one finds

$$X_{ij} = \nabla_{\hat{i}} \nabla_{\hat{j}} Z = J_i^k J_j^l \nabla_k \nabla_l Z \quad J_i^j : \text{complex structure}$$

and  $E_k^k = -Y_2$

so  $\square A = \frac{\beta}{1728} Y_2$

$$R_{ij} = \frac{\beta}{1152} \left( \nabla_{\hat{i}} \nabla_{\hat{j}} Z + \nabla_i \nabla_j Z + 2E_{ij} + \frac{1}{4} Y_2 g_{ij} \right)$$

terms expected  
from  $CY_3$  case

terms arising  
from  $Y_2$

- ◆ These corrections have the effect of making the Ricci tensor non-vanishing, and even remove the Kähler property of the metric. Nonetheless, the manifold remains special, as we shall see.



# Gravitational sourcing of 4-form flux

- ◆ The corrected 3-form field equation is

$$d * G + \frac{1}{2} G \wedge G + (2\pi)^4 \beta X_8 = 0$$

- for initial purely gravitational backgrounds with  $c_4 = 0$ , this forces 4-form flux to turn on at order  $\beta$

- ◆ Let  $\hat{G}_{[4]} = G_{[3]} \wedge d\tau + G_{[4]}$

- for  $c_4 = 0$ , assume  $G_{[4]} = 0$

- then  $d * G_{[3]} = (2\pi)^4 \beta X_8$  D=10 Hodge dual here

- in turn, write  $G_{[3]} = \frac{3}{4} J \wedge dA + \tilde{G}_{[3]}$   $J^{jk} \tilde{G}_{ijk} = 0$

- ◆ Then the Einstein equation becomes

$$R_{ij} = \frac{3}{8} (\nabla_i \nabla_j A + \nabla_{\hat{i}} \nabla_{\hat{j}} A) + \frac{\beta}{1152} (\nabla_i \nabla_j Z + \nabla_{\hat{i}} \nabla_{\hat{j}} Z) - \frac{1}{2} \nabla^k \tilde{G}_{i\hat{j}k}$$

- ◆ The gravitational sourcing of 4-form flux is accompanied by changes to the Killing spinor and to the complex structure.

- ◆ Killing spinor equation:  $\hat{D}_m \eta = 0$  becomes deformed, requiring a brane-like warp factor  $\hat{\eta} = e^{\frac{1}{2}A} \eta$  and

$$D_i \eta = \nabla_i \eta + i(\nabla_i h) \eta + \frac{i}{8} \tilde{G}_{ijk} \gamma^{jk} \eta = 0$$

$$h = \frac{3}{16} A + \frac{\beta}{2304} Z \quad \gamma_{11} \eta = -\eta \quad \tilde{G}_{ijk} \gamma^{ijk} \eta = 0$$

- ◆ The deformed Killing spinor leads to a deformed complex structure  $J_{ij} = -i \bar{\eta} \gamma_{ij} \eta$

$$\nabla_j J_i^j = \frac{1}{2} \tilde{G}_{ij}^k - \frac{1}{2} \tilde{G}_{\hat{i}j}^{\hat{k}} \neq 0$$

so the deformed space is no longer Kähler

- ◆ Despite the loss of Kähler structure, the Nijenhuis tensor  $N_{ij}{}^k = \partial_{[i}J_{j]}{}^k - J_i{}^l J_j{}^m \partial_{[m}J_{l]}{}^k$  still vanishes, so the deformed space is still a complex manifold.
- ◆ It no longer has  $SU(5)$  holonomy, but one may still define a generalized holonomy for the Killing spinor operator  $D_i$ . The generalized transverse structure group is  $SL(16, \mathbb{C})$ . The decomposition of the deformed Killing spinor under the generalized holonomy still contains singlets, showing that supersymmetry remains unbroken.

# Deformed special holonomy

- ◆ The effect of the  $\alpha'^3$  corrections is to destroy the original special holonomy, giving a general complex D=10 manifold.
- ◆ Nonetheless, the specific structure of the  $\alpha'^3$  corrections is such as to *permit* the corrected Einstein equation to arise as the integrability condition for an  $\alpha'^3$  corrected Killing spinor equation.
- ◆ This fits into a general pattern that obtains also for 7-manifolds of initially  $G_2$  and 8-manifolds of initially  $Spin_7$  holonomies.

- ◆ In all cases (including the  $D \leq 8$  Kähler cases where the effect of the corrections is simply to include an extra  $U(1)$  factor in the holonomy), supersymmetry can be preserved providing the Killing spinor equation acquires its own  $\alpha'^3$  correction, e.g. for the  $D \leq 8$  cases

$$D_i \eta = (\nabla_i + \xi(\alpha')^3 Q_i) \eta = 0$$

where

Candelas, Freeman, Pope, Sohnius & K.S.S

$$Q_i = -\frac{3}{4} (\nabla^j R_{ikm_1m_2}) R_{jlm_3m_4} R^{kl}_{m_5m_6} \Gamma^{m_1 \dots m_6}$$

- ◆ In the various cases of initially special holonomy, this can be rewritten in ways that more directly yield the corrected Einstein equation as integrability condition,

e.g. in the  $G_2$  case  $Q_i = -\frac{1}{2} i c_{ijk} \nabla^j Z^{kl} \tilde{\Gamma}_l$

while in the  $Spin_7$  case  $Q_i = \frac{1}{4} c_{ijkl} \nabla^j Z^{klmn} \Gamma_{mn}$

# Generalized Structure groups and holonomy

- ◆ Although the ordinary Riemannian holonomy becomes generic for the corrected internal spaces, the supersymmetry preservation can still be understood on group theoretical grounds, using the notion of generalized holonomy.
- ◆ Consider the transverse groups generated by the generic Gamma matrix combinations present in the corrected Killing operator ( $\Gamma_{[2]}$ ,  $\Gamma_{[6]}$  and their closure), restricting attention to the  $D$  “transverse” dimensions only:
  - $D=7$   $SO(8)$
  - $D=8$   $SO(8)_+ \times SO(8)_-$
  - $D=10$   $SL(16, \mathbb{C})$

- ◆ Within these generalized transverse structure groups, the generalized holonomy is the group actually generated by the operators present in the corrected Killing spinor operator for a given space. Under decomposition into representations of these groups, the spinor representation contains a singlet, indicating continued supersymmetry preservation:

- $D=7 \quad SO(8) \rightarrow SO(7) \quad (\text{corrected } G_2)$   
 $8_{\pm} \rightarrow 7 \oplus 1$

- $D=8 \quad SO(8)_+ \otimes SO(8)_- \rightarrow SO(8)_+ \otimes (\text{Spin}_7)_- \quad (\text{corrected } \text{Spin}_7)$   
 $(8, 1) \oplus (1, 8) \rightarrow (8, 1) \oplus (1, 7) \oplus (1, 1)$

- $D=10 \quad SL(16, \mathbb{C}) \rightarrow [U(1) \times SL(5, \mathbb{C}) \times SL(5, \mathbb{C})] \times [\mathbb{C}_1^{(10,1)} \oplus \mathbb{C}_3^{(10,5)}]$

16 rep once again decomposes including a singlet